LOOK BEFORE YOU LEAP

Using the Solar Gravitational Lens to explore exoplanets

The scale of the problem



Extrasolar scales

• Distance to the nearest star is almost 300,000 AU

- That is almost 100 million times the famous "one giant leap"
- If the Earth is a grape and the Moon is a peppercorn a foot away...
- ... the nearest star would be more than halfway around the Earth.
- It makes sense to look before we take such a giant leap

What does it entail to look?

- Fundamentally, three issues:
 - Resolution: We want to be able to see details of a distant world. Diffraction
 - Brightness: We want enough light to form an image. Photons are scarce
 - Noise: There are many contaminating sources of light. Noise can get amplified

The tyranny of the diffraction limit

- Angular resolution is roughly proportional to λ/d
- Resolving features of 10 km at 10 light years implies d ~ 10,000 km
- Large baselines and interferometry may help but there's noise
 - Light contamination from host star
 - Exozodiacal noise
 - Other techniques (e.g., rotational deconvolution) have limitations

The Sun to the rescue!

The Sun bends rays of light and acts as an imperfect lens

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The focal region

• Parallel rays of light are deflected, meeting at 550+ AU



Projection vs. view

The Sun projects a km-scale blurry image of a distant object
A telescope in the focal region only sees a thin Einstein-ring

1.3 km |---

Images must be reconstructed

- Multiple observations needed
- Each observation captures 1 "pixel"
- What is imaged is a moving target
- Long exposure times are required
- Precise stationkeeping, location tracking essential

Why not go there today?

Target must be known in advance to pick focal region

- Imaging has many challenges
- Distance is enormous

The imaging challenge: signal-to-noise

- From any one vantage point in the image plane, all we see is a very faint Einstein ring through the very bright solar corona
- Corona is 10⁴ 10⁵ times brighter than the Einstein ring
 - Even if the corona background is removed, we have to deal with shot noise due to the very low signal photon count
 - We must also block light from the Sun
 - There will be light from the host star, brightening the Einstein ring in specific places
 - Light from interlopers must be known

The imaging challenge: collecting light

- Low SNR and stochastic (shot) noise has one solution: increased integration times
- Prolongs mission duration
 - Represents added navigational challenge
 - Must deal with a temporally changing target

The imaging challenge: imperfect lens

- The monopole SGL is a lens with significant spherical aberration
- Quadrupole contributes significant astigmatism
- Reconstruction (deconvolution) is possible but it substantially amplifies noise
- Without deconvolution, a less noisy but blurry image results
 - Trade-off between noise, resolution and sharpness must be considered

The imaging challenge: moving target

- An Earth-like planet rotates, changes significantly over the scale of minutes
- Its illumination changes with the seasons
- Its appearance changes: short-term stochastic changes (weather) and periodic changes (e.g., ice cover, vegetation)

Recovering a moving target



Set Set



Load dataset Save dataset 5856 observations so far. The Julian date is 2460088 and 23 hours.

unter:	function stepCount(){ return 500;}	
isor:	function stepDivisor(){ return 2000;}	

Model height: 44, width: 133 Residual: 0.301, SNR: 1.71









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How to get there?



8th Interstellar Symposium, Montreal, July 10, 2023

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Navigation is a serious challenge

Host star is visible throughout the cruise

• Focal region wobbles by as much as a million km



Navigating from 400,000 to 4,000 km

Light amplification by O(100)



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Zeroing in on the host star

• From 1,000 km to the exostar optical axis





Fine-grained navigation

- Exoplanet image located O(10,000 km) from host star image in the image plane
- Approximate direction known
- Approaching the exoplanet image is a repeat of the host star approach but on a smaller and fainter scale, on the host star light background
- Exoplanet image size is of O(1 km)

Need for a local reference frame

- Projected image motion is known but it is noninertial
- Absolute navigation at that accuracy is unrealistic
- Once the exoplanet image is located, two or more telescopes could establish its precise boundary
- These telescopes establish and maintain a local reference frame
- Observing telescope can move from pixel to pixel on a meter-scale grid with respect to this frame
- Primary observable: Total amount of light collected from the Einsteinring at each pixel location

The means to get there

• Sundivers!





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Technical challenges

- Large spacecraft vs. in-flight assembly
- Systems to survive 25-30 years cruise plus up to 10 years science
- Autonomous constellation (required for navigation, corona light suppression)
- Coronagraph (possible use of discarded sails as external sunshades?)
- Power (40+ years)
- Communication (650 900 AU)

Thank you

• Questions?